

DEFINING ENTROPY BOUNDS *

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Abstract

Bekenstein's conjectured entropy bound for a system of linear size R and energy E , $S \leq 2\pi ER$, has counterexamples for many of the ways in which the "system," R , E , and S may be defined. Here new ways are proposed to define these quantities for arbitrary nongravitational quantum field theories in flat spacetime, such as defining R as the smallest radius outside of which only vacuum expectation values occur. Difficulties of extending these definitions to gravitational quantum and semiclassical theories are noted.

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1 Introduction

Bekenstein has conjectured [1] that the entropy S of a system confined to radius R or less and energy E or less would obey the inequality (using units $\hbar = c = k_{\text{Boltzmann}} = 1$)

$$S \leq 2\pi ER. \quad (1)$$

He and colleagues have supported this conjecture with many arguments and examples [1-18]. However, many counterarguments and counterexamples have also been noted [19-34]. Whether the conjectured bound (1) holds or not depends on what systems are considered and how R , E , and S are defined.

Perhaps the simplest procedure [8, 10] would be to just consider quantum fields inside some bounded region within a sphere of radius R and put boundary conditions on the fields at the boundary of the region. However, this procedure leads to a large number of counterexamples to Bekenstein's conjectured bound. For example [20], the Casimir effect can make $E < 0$ for certain states of quantum fields confined within a certain regions of radius $\leq R$, violating the bound. If states with $E < 0$ are excluded by definition, one can still consider a mixed state with arbitrarily small positive E that violates the bound. Even if E is redefined to be the nonnegative energy excess over that of the ground state [8, 10], one can violate the bound by a mixed state that is almost entirely the ground state and a tiny incoherent mixture of excited states, at least if the entropy is defined to be $S = -\text{tr} \rho \ln \rho$ [22]. If S is instead defined to be $S = \ln n$ for a mixture of n orthogonal pure states (which would agree with $S = -\text{tr} \rho \ln \rho$ if the mixture had equal probabilities $1/n$ for each of those n pure states), then one can violate the bound by an equal mixture of the ground state and the first excited state of certain interacting fields with certain boundary conditions that have the two lowest states nearly degenerate in energy (separated by exponentially small tunneling effects) [20, 32, 33]. If interacting fields are excluded from the definition of allowable systems, one can get a violation by considering a sufficiently large number N of identical free fields, giving $n = N$ degenerate first excited states of finite energy but sufficiently large entropy $S = \ln n = \ln N$ to violate (1) [20, 21]. And even for a single free electromagnetic field, $S = \ln n$ can exceed $2\pi ER$ by an arbitrarily large factor by using boundary conditions corresponding to an arbitrarily large number of parallel perfectly conducting plates within the region of radius R [32], or by using boundary conditions corresponding to an arbitrarily long coaxial cable loop coiled up within the region [33].

However, other than in his papers with Schiffer [8, 10], Bekenstein has generally advocating taking E to be the total energy of a complete system [1, 3, 4, 6, 15, 18]. This would disallow using just the energy of fields within a bounded region with boundary conditions, since that would ignore the energy of whatever it is that is

providing the boundary conditions. Therefore, all of the counterexamples mentioned above would be excluded by this restriction. However, then the problem is to define what one means by the radius R of the system. In the weakly gravitating case (essentially quantum fields in flat Minkowski spacetime) that we shall focus on here, Bekenstein takes R to mean the radius of a sphere which circumscribes the system, which leaves the problem of what it means for a sphere to circumscribe the complete system.

In quantum field theory in Minkowski spacetime, the complete system is the quantum state of the fields. Since the quantum fields extend all the way out to radial infinity, a sphere circumscribing the entire system would have to be at $R = \infty$, which makes the Bekenstein bound true (at least for states of positive energy and finite entropy) but trivial. To get a nontrivial bound, one needs to suppose that a sphere of finite R can circumscribe the system. For example, one might try to say that the sphere encloses all of the excitations of the fields from the vacuum. However, it is also hard to get this to occur for a finite R . For example, the wavefunction for any single particle state that is a superposition of energy eigenstates of bounded energy will not vanish outside any finite radius R , since a one-particle wavefunction that does vanish outside a finite region must be a superposition of arbitrarily large momentum components, which will have unbounded energy. Even if one looks at a composite system, such as a hydrogen atom, and ignores the fact that its center of mass will have amplitudes to be outside any finite sphere if it is made of purely bounded energy components, the wavefunction for the relative position of the electron and proton does not drop identically to zero outside any finite separation distance for states that are superpositions of energy eigenstates of bounded energy. In particular, even if one fixed the center of mass of a hydrogen atom in its ground state and ignored the infinite energy from the resulting infinite uncertainty of the center of mass momentum, the density matrix for the electron position would decay only exponentially with distance from the center of mass and never go to zero outside any sphere of finite radius R .

Therefore, it is problematic to define the radius R of a sphere circumscribing a complete system in any quantum field theory. This issue has not been addressed by Bekenstein and his collaborators, but without such a definition, there is no nontrivial formulation of the conjectured bound (1) for complete systems, but only its trivial truth for any complete system with positive energy and finite entropy that can only be circumscribed by the sphere enclosing all of space, $R = \infty$.

Here new ways are proposed to define systems and their radii R , energies E , and entropies S , so that for each, there is a bound on S for a given system as a function of finite R and E . These bounds will not have the form of Bekenstein's conjectured inequality (1), though in some cases they may obey that inequality.

2 Vacuum-Outside-R States

The main new element of the present paper is a proposal is to define a system of radius R (in flat spacetime for the present) not by imposing boundary conditions on the field itself, but by imposing conditions on the quantum state of the field so that outside a closed ball of radius R the quantum state is indistinguishable from the vacuum at some time. Such a state will be called a vacuum-outside- R state. (For simplicity, set this time to be $t = 0$, and take the closed ball, say B , to be the region $r \leq R$ on the $t = 0$ hypersurface, where r is the standard radial polar coordinate giving the proper distance from the coordinate origin on that hypersurface.) In other words, a vacuum-outside- R state of the system, say as expressed by its density matrix ρ , is such that the expectation value of any operator O which is completely confined to the region $r > R$ when written in terms of field and conjugate operators at $t = 0$, is precisely the same as the expectation value of the same operator in the vacuum state $|0\rangle\langle 0|$,

$$\text{tr}(O\rho) = \langle 0|O|0\rangle. \quad (2)$$

In particular, all the n -point functions for the field and for its conjugate momentum in the state ρ are the same as in the vacuum state, if all of the n points are outside the ball of radius R and on the hypersurface $t = 0$. Of course, the n -point functions need not be the same as their vacuum values if some or all of the points are inside the ball.

If operators confined to the three-dimensional region $r > R$ and $t = 0$ (say C , to give a name to this achronal spacelike surface, the $t = 0$ hypersurface with the central closed ball B , $r \leq R$, excluded) have the same expectation value in the vacuum-outside- R state as in the vacuum state, the same will be true in any quantum field theory that I shall call “strongly causal” for all operators confined to the Cauchy development or domain of dependence [35] of C , the larger four-dimensional region $r > R + |t|$ (say D) that is the set of all points in the Minkowski spacetime such that every inextendible (endless) causal, or non-spacelike (everywhere timelike or lightlike), curve through such a point intersects the partial Cauchy surface C . Just as solutions of hyperbolic wave equations in D are determined by the data on C , so the part of the quantum state of a strongly causal field in D , as represented by the expectation values of operators confined to D , is determined by the part of the quantum state in C , as represented by the expectation values of operators confined to C . (For some interacting quantum field theories, the expectation values of operators confined to the three-dimensional spacelike surface C may be too ill-defined for these theories to be “strongly causal” in my sense, but a wider class of these theories may be “weakly causal” in the sense that sufficiently many operators smeared over, but confined to, an arbitrarily thin-in-time four-dimensional slab, say

E , containing C within D , have well-defined expectation values that determine the expectation values of all operators smeared over, but confined to, any part of D .)

Henceforth I shall restrict attention to strongly causal and weakly causal quantum field theories, calling them simply causal quantum field theories for short. I shall also assume, until discussing gravitational theories later, that any quantum field theory under consideration is a nongravitational Lorentz-invariant quantum field theory in Minkowski spacetime, and that it has a unique pure state of lowest Minkowski energy $E = 0$ (the expectation value of the Hamiltonian H that generates translations in the time coordinate t in some Lorentz frame, with the arbitrary constant in the Hamiltonian being adjusted to give the lowest energy state zero energy).

Therefore, for such a causal nongravitational quantum field theory in Minkowski spacetime, I shall propose that the radius R be defined so that all of the operators constructed from field and conjugate momentum operators smeared over regions confined to the region D , $r > R + |t|$ in some Lorentz frame, have in the particular quantum state being considered (a vacuum-outside- R state) the same expectation values that they have in the vacuum state for that quantum field theory. The energy E of the state ρ can then be simply defined to be the expectation value,

$$E \equiv \text{tr}(H\rho), \quad (3)$$

of the Hamiltonian H that generates time translations in the same Lorentz frame. Because the energy E has been defined to have the minimum value of zero for the unique pure vacuum state, there is no problem here with negative Casimir energies. In other words, the energy is that of the complete system over all of Minkowski spacetime.

Obviously we would also like a definition of the entropy S that has a minimum value of zero, which it should attain for the pure vacuum state. One simple definition is the von Neumann entropy,

$$S = S_{\text{vN}} \equiv -\text{tr} \rho \ln \rho, \quad (4)$$

using the density matrix ρ for the full state of the quantum field, over the entire Minkowski spacetime.

3 Entropy Bounds for Vacuum-Outside- R States

Now we may conjecture that for any vacuum-outside- R state of any particular causal nongravitational quantum field theory in Minkowski spacetime, one which has the vacuum expectation values in the region D , $r > R + |t|$ (the region causally disconnected from the ball $r \leq R$ at $t = 0$), the von Neumann entropy is bounded above

by some function σ_{vN} (depending on the quantum field theory in question) of the radius R and energy E :

$$S_{\text{vN}} \leq \sigma_{\text{vN}}(R, E). \quad (5)$$

Define this function $\sigma_{\text{vN}}(R, E)$ to be the least upper bound on the von Neumann entropy of any state which is vacuum outside the radius R and which has energy E .

In the case of a scale-invariant quantum field, such as a free massless field, or say a massless scalar field ϕ with a $\lambda\phi^4$ self-coupling potential, the least upper bound function $\sigma_{\text{vN}}(R, E)$ will actually be a function of the single dimensionless variable

$$x \equiv 2\pi RE, \quad (6)$$

say

$$\sigma_{\text{vN}}(R, E) = \sigma_{\text{N}}(x). \quad (7)$$

Bekenstein's conjectured entropy bound (1), if R , E , and S were defined as done herein, would be $\sigma_{\text{vN}}(R, E) \leq x$, whether or not the quantum field theory is scale invariant, or

$$B_{\text{vN}}(R, E) \equiv \frac{\sigma_{\text{vN}}(R, E)}{x} \equiv \frac{\sigma_{\text{vN}}(R, E)}{2\pi RE} \leq 1. \quad (8)$$

If the quantum field theory is scale invariant, we can define

$$B_{\text{N}}(x) \equiv \frac{\sigma_{\text{N}}(x)}{x}, \quad (9)$$

which should also be less than or equal to unity if Bekenstein's bound applies.

For a set of one or more free massless fields and vacuum-outside- R states with $x \gg 1$, one would expect that the highest entropy would be given by a mixed state that at $t = 0$ is approximately a high-temperature ($RT \gg 1$) thermal radiation state for $r < R$, surrounded by vacuum for $r > R$. A high-temperature thermal radiation state has an energy density for massless fields of approximately $a_r T^4$, and hence an entropy density $(4/3)a_r T^3$, where

$$a_r = \frac{\pi^2}{30} \left(n_b + \frac{7}{8} n_f \right) \quad (10)$$

is the radiation constant for n_b independent bosonic degrees of freedom for each momentum (e.g., n_b different spin or helicity states) and for n_f fermionic degrees of freedom. Therefore, in this case with $x \gg 1$,

$$B_{\text{vN}}(R, E) = \frac{\sigma_{\text{N}}(x)}{x} \approx \left(\frac{2^7 a_r}{3^5 \pi^2 x} \right)^{\frac{1}{4}} = \left[\frac{2^6}{3^6 5 x} \left(n_b + \frac{7}{8} n_f \right) \right]^{\frac{1}{4}}, \quad (11)$$

which is indeed less than 1, thus obeying Bekenstein's conjectured bound, for

$$x \geq \frac{2^7 a_r}{3^5 \pi^2} = \frac{2^6}{3^6 5} \left(n_b + \frac{7}{8} n_f \right) = \frac{64 n_b + 56 n_f}{3645}, \quad (12)$$

if x is also large enough that Eq. (11) is a good approximation. Thus one would expect that Bekenstein's conjectured bound, using the definitions above for R , E , and S , holds for a fixed set of free massless quantum fields at sufficiently large $x \equiv 2\pi RE$.

On the other hand, the definitions above for R , E , and S still permit Bekenstein's conjectured bound applied to them to be violated for sufficiently small x , as we can see by the following construction:

A way to construct vacuum-outside- R states, quantum states of a free quantum field theory that have vacuum expectation values in the region D , $r > R + |t|$, is to apply to the vacuum state unitary operators constructed from fields and/or conjugate momenta smeared within the region $r < R$ at $t = 0$. In particular, if h is an hermitian operator constructed from fields and/or conjugate momenta smeared within $r < R$ at $t = 0$, then $U = e^{ih}$ is such a unitary operator, and

$$|\psi\rangle = U|0\rangle = e^{ih}|0\rangle \quad (13)$$

is a pure quantum state that has precisely the vacuum expectation values in the region D . This result can be seen formally from the fact that any operator O confined to the region D (the four-dimensional region $r > R + |t|$) that is causally disconnected from the ball B (the three-dimensional region $r \leq R$ on the $t = 0$ hypersurface) commutes with the operators h and U that are confined to that hypersurface, $[O, h] = [O, U] = 0$, so

$$\text{tr}(O\rho) = \langle\psi|O|\psi\rangle = \langle 0|U^{-1}OU|0\rangle = \langle 0|U^{-1}UO|0\rangle = \langle 0|O|0\rangle. \quad (14)$$

If $\{h_i\}$ is a set of hermitian operators that each are confined to the ball B (i.e., are constructed from fields and momenta that are smeared only over that region), and if $\{q_i\}$ is a set of positive numbers that sum to unity, then

$$\rho = \sum_i q_i e^{ih_i}|0\rangle\langle 0|e^{-ih_i} \quad (15)$$

is a more general vacuum-outside- R state, since this density matrix gives vacuum expectation values, $\text{tr}(O\rho) = \langle 0|O|0\rangle$, for any operator O confined to the region D that is causally disconnected from B (i.e., having no causal curves, either timelike or lightlike, intersecting both the ball B of $r \leq R$ at $t = 0$ and the region D with $r > R + |t|$).

Here let us consider the simple example in which i takes only the two values 1 and 2, and $h_1 = 0$ and $h_2 = h$. Let $q_1 = 1 - q$ and $q_2 = q$, and let

$$e^{ih}|0\rangle = U|0\rangle = |\psi\rangle = c|0\rangle + s|1\rangle \quad (16)$$

in terms of a decomposition of $|\psi\rangle$ into the two orthonormal states $|0\rangle$ and $|1\rangle = (|\psi\rangle - \langle 0|\psi\rangle|0\rangle)/\sqrt{1 - |\langle 0|\psi\rangle|^2}$, so

$$c = \langle 0|\psi\rangle = \langle 0|U|0\rangle = \langle 0|e^{ih}|0\rangle, \quad (17)$$

$$s = \sqrt{1 - |\langle 0|U|0\rangle|^2} = \sqrt{1 - |c|^2}. \quad (18)$$

Note that $|1\rangle\langle 1|$ by itself is not generically a vacuum-outside- R state.

Now Eq. (15) gives the density matrix as

$$\begin{aligned} \rho &= (1 - q)|0\rangle\langle 0| + q|\psi\rangle\langle\psi| \\ &= (1 - qs^2)|0\rangle\langle 0| + qcs|0\rangle\langle 1| + q\bar{c}s|1\rangle\langle 0| + qs^2|1\rangle\langle 1|, \end{aligned} \quad (19)$$

a density matrix in the two-dimensional space of pure states spanned by the two orthonormal pure states $|0\rangle$ and $|1\rangle$. The two eigenvalues of this density matrix are, say, p and $1 - p$ (since their sum is $\text{tr}\rho = 1$), with product

$$\begin{aligned} y \equiv p(1 - p) &= \frac{1}{2}\{[p + (1 - p)]^2 - [p^2 + (1 - p)^2]\} = \frac{1}{2}\{[\text{tr}(\rho)]^2 - [\text{tr}(\rho^2)]\} \\ &= q(1 - q)(1 - |\langle 0|U|0\rangle|^2) = q(1 - q)s^2. \end{aligned} \quad (20)$$

The expectation value of the energy of this mixed state is, since I have assumed $H|0\rangle = 0$,

$$E = \text{tr}(H\rho) = q\langle\psi|H|\psi\rangle = q\langle 0|U^{-1}HU|0\rangle. \quad (21)$$

Then

$$x \equiv 2\pi RE = 2\pi Rq\langle 0|U^{-1}HU|0\rangle. \quad (22)$$

The von Neumann entropy of this mixed state is

$$\begin{aligned} S_{\text{vN}}(y) &= -\text{tr}(\rho \ln \rho) = -p \ln p - (1 - p) \ln (1 - p) \\ &= \frac{y}{\frac{1}{2}(1 + \sqrt{1 - 4y})} \ln \frac{1}{y} + \sqrt{1 - 4y} \ln \frac{1}{\frac{1}{2}(1 + \sqrt{1 - 4y})} \\ &\approx y[(1 + y + 2y^2) \ln \frac{1}{y} + (1 - \frac{1}{2}y - \frac{5}{3}y^2)], \end{aligned} \quad (23)$$

a monotonically increasing function of $y \equiv q(1 - q)s^2 \leq 1/4$, where the last approximate equality of Eq. (23) applies for $y \ll 1$.

As q and/or h is reduced toward zero, x , y , and S also decrease toward zero, but whereas x and y asymptotically decrease linearly with q , the dominant term of S has an extra logarithmic factor that grows with the reduction of y , so the ratio,

$$B \equiv \frac{S_{\text{vN}}}{x} \equiv \frac{S_{\text{vN}}}{2\pi RE} \approx \frac{y}{x} \left(\ln \frac{1}{y} + 1 \right) \quad (24)$$

when $y \ll 1$, increases without limit as y is reduced toward zero. Therefore, when y is made sufficiently small (e.g., by making q sufficiently small), Bekenstein's conjectured bound for the definition of R , E , and S used here is violated.

4 Free Quantum Field Theory Examples

Let us consider a specific example for the Hermitian operator h that is constructed from field operators confined to the ball B of radius R on an initial flat hypersurface of Minkowski spacetime. Take the quantum field theory to be that of a single massless scalar field operator ϕ . Consider the smeared linear Hermitian field operator

$$\chi = \int d^3x [F(\mathbf{x}) \phi(t=0, \mathbf{x}) + G(\mathbf{x}) \dot{\phi}(t=0, \mathbf{x})], \quad (25)$$

where $F(\mathbf{x})$ and $G(\mathbf{x})$ are real functions of the spatial location \mathbf{x} that are zero for $|\mathbf{x}| > R$, so that χ is made up of operators confined to the ball B , $r \leq R$ at $t = 0$. Then for real parameters α and β , let

$$h = \alpha\chi + \beta\chi^2, \quad (26)$$

which is thus also an Hermitian operator confined to the ball B .

Then by expanding out $\phi(\mathbf{x})$ and H in terms of creation and annihilation operators, one can show, after a certain amount of algebra that will not be repeated here, that

$$c \equiv \langle 0|\psi \rangle \equiv \langle 0|U|0 \rangle \equiv \langle 0|e^{ih}|0 \rangle = (1 - 2i\beta X)^{-1/2} \exp\left(\frac{-\frac{1}{2}\alpha^2 X}{1 - 2i\beta X}\right), \quad (27)$$

$$s^2 \equiv 1 - |\langle 0|U|0 \rangle|^2 \equiv 1 - |c|^2 = 1 - (1 + 4\beta^2 X^2)^{-1/2} \exp\left(\frac{-\alpha^2 X}{1 + 4\beta^2 X^2}\right), \quad (28)$$

and

$$\langle \psi|H|\psi \rangle \equiv \langle 0|U^{-1}HU|0 \rangle \equiv \langle 0|e^{-ih}He^{ih}|0 \rangle = \alpha^2 Y + 4\beta^2 XY = (\alpha^2 X + 4\beta^2 X^2)Z, \quad (29)$$

where

$$X \equiv \langle 0|\chi^2|0 \rangle = \int d^3x d^3y \frac{F(\mathbf{x})F(\mathbf{y}) + \nabla G(\mathbf{x}) \cdot \nabla G(\mathbf{y})}{4\pi^2 |\mathbf{x} - \mathbf{y}|^2} \quad (30)$$

and

$$Y \equiv XZ \equiv \langle 0|\chi H \chi|0 \rangle = \frac{1}{2} \int d^3x [|F(\mathbf{x})|^2 + |\nabla G(\mathbf{x})|^2]. \quad (31)$$

Incidentally, I have not included individual higher powers of χ in h , because then expanding e^{ih} into a power series in χ and taking the expectation values gives divergent series when one uses the key intermediate results

$$\langle 0|\chi^m|0 \rangle = \begin{cases} (m-1)!! X^{m/2} & m \text{ even} \\ 0 & m \text{ odd} \end{cases}, \quad (32)$$

and

$$\langle 0|\chi^m H \chi^n|0 \rangle = \begin{cases} mn(m+n-3)!! X^{(m+n)/2} Z & m+n \text{ even} \\ 0 & m+n \text{ odd} \end{cases}. \quad (33)$$

(The divergences arise from the rapid growth of the double factorials with their arguments. These double factorials arise from the counting of the number of pairings of the creation and annihilation operators in the powers of the χ 's and in the Hamiltonian H for the massless scalar field ϕ .)

If one takes $F(\mathbf{x})$ and $G(\mathbf{x})$ to be spherically symmetric, say

$$F(\mathbf{x}) = R^{-2} f\left(\frac{|\mathbf{x}|}{R}\right) \equiv R^{-2} f(u) \quad (34)$$

and

$$G(\mathbf{x}) = R^{-1} g\left(\frac{|\mathbf{x}|}{R}\right) \equiv R^{-1} g(u) \quad (35)$$

with f and g being dimensionless functions of the dimensionless radial variable (hereafter to be called u or v) that vanish when the latter variable is greater than unity (corresponding to points outside the sphere of radius R), then

$$X = \int_0^1 du \int_0^1 dv \left\{ 2uv \ln \left| \frac{u+v}{u-v} \right| f(u)f(v) + \left[(u^2 + v^2) \ln \left| \frac{u+v}{u-v} \right| - 2uv \right] g'(u)g'(v) \right\} \quad (36)$$

and

$$Y \equiv XZ = \frac{2\pi}{R} \int_0^1 u^2 du [f^2(u) + g'^2(u)], \quad (37)$$

where the prime on the function g denotes a derivative with respect to the argument (the dimensionless radius $u \equiv |\mathbf{x}|/R$ or $v \equiv |\mathbf{y}|/R$).

Now, if we take a density matrix of the form (19), let us try to maximize the product of the two nonzero eigenvalues of the density matrix,

$$y \equiv p(1-p) = q(1-q)s^2 = q(1-q) \left[1 - (1 + 4\beta^2 X^2)^{-1/2} \exp \left(\frac{-\alpha^2 X}{1 + 4\beta^2 X^2} \right) \right], \quad (38)$$

and hence maximize $S_{\text{vN}}(y)$ given by Eq. (23), for fixed

$$x \equiv 2\pi RE = 2\pi Rq <0|U^{-1}HU|0> = 2\pi RZq(\alpha^2 X + 4\beta^2 X^2). \quad (39)$$

Note that for fixed $RZ = RY/X$, the three quantities α , β (the coefficients of χ and of χ^2 in the hermitian operator $h = \alpha\chi + \beta\chi^2$), and $X \equiv <0|\chi^2|0>$ enter into this x and y only in the two nonnegative combinations $a \equiv \alpha^2 X$ and $b \equiv 4\beta^2 X^2$, and x depends only on $q(a+b)$. Then it is easy to see that for fixed q and fixed $a+b$, y decreases monotonically with b , so to maximize y and $S_{\text{vN}}(y)$ for fixed x , we should set $\beta = 0$ in order to get $b = 0$, $h = \alpha\chi$,

$$a \equiv \alpha^2 X = <0|h^2|0>, \quad (40)$$

$$x = 2\pi RZqa, \quad (41)$$

and

$$y = q(1 - q)(1 - e^{-a}). \quad (42)$$

Next, in our attempt to maximize y as a function of x , we may continue to hold

$$\gamma \equiv 2\pi RZ \quad (43)$$

fixed and hence maximize y for fixed $z \equiv x/\gamma = aq$. Then one can easily calculate that $y = q(1 - q)(1 - e^{-z/q})$ has its maximum at fixed z when

$$q = \frac{e^a - 1 - a}{2e^a - 2 - a} \approx \frac{1}{2}a(1 - \frac{2}{3}a), \quad (44)$$

giving

$$x = \gamma a q = \frac{\gamma a(e^a - 1 - a)}{2e^a - 2 - a} \approx \frac{1}{2}\gamma a^2(1 - \frac{2}{3}a) \quad (45)$$

and

$$y = [1 - (1 + a)e^{-a}] \left(\frac{e^a - 1}{2e^a - 2 - a} \right)^2 \approx \frac{1}{2}a^2(1 - \frac{5}{3}a) \approx \frac{x}{\gamma}, \quad (46)$$

where all the approximate equalities apply for $a \ll 1$. For fixed γ , a is given implicitly as a function of x by Eq. (45), and then inserting Eq. (46) for y into Eq. (23) for $S_{\text{vN}}(y)$ gives the entropy (so far maximized over β and q) of the density matrix (19) explicitly as a function of $a = \langle 0|h^2|0 \rangle$ and hence implicitly as a function of x . In fact, for $x \ll \gamma$, we get the asymptotic relation

$$S_{\text{vN}}(y) \sim \frac{x}{\gamma} \left(\ln \frac{\gamma}{x} + 1 \right), \quad (47)$$

which of course exceeds x for sufficiently small x .

As the final step in the maximization of the von Neumann entropy $S_{\text{vN}}(y)$ of a density matrix of the particular form (19) for fixed $x \equiv 2\pi RE = \gamma z = \gamma a q$, we note that maximizing y (and hence $S_{\text{vN}}(y)$) for fixed x is equivalent to minimizing x for fixed y . Therefore, we need to minimize $\gamma \equiv 2\pi RZ \equiv 2\pi RY/X$. By looking at Eqs. (30) and (31) for X and Y , we see that the ratio $Z \equiv Y/X$ is invariant under any constant rescaling of the functions $F(\mathbf{x})$ and $G(\mathbf{x})$ that appear as smearing functions for $\phi(t=0, \mathbf{x})$ and $\dot{\phi}(t=0, \mathbf{x})$ in the defining Eq. (25) for the linear hermitian field operators χ and $h = \alpha\chi$ [now that we have set $\beta = 0$ to drop the nonlinear term for h in Eq. (26)]. The quantity γ is also invariant under a rescaling of the radius R if $F(\mathbf{x})$ and $G(\mathbf{x})$ depend only on \mathbf{x}/R and on some overall constant factor that can depend on R .

Minimizing $\gamma = 2\pi RY/X$ is thus equivalent to maximizing X for fixed Y , which by Eq. (31) is half the integral of the sum of the squares of $F(\mathbf{x})$ and of the gradient of $G(\mathbf{x})$. Because the double integral (30) for X is also quadratic in $F(\mathbf{x})$ and

in $\nabla G(\mathbf{x})$ but has a positive-definite nonlocal kernel, maximizing X at fixed Y is best done with fairly smooth functions $F(\mathbf{x})$ and $G(\mathbf{x})$. In particular, if $F(\mathbf{x})$ is expanded in spherical harmonics, one can readily see that the maximum is obtained by keeping only the spherically symmetric ($\ell = 0$) terms. Also, since it is the dot product of $\nabla G(\mathbf{x})$ and $\nabla G(\mathbf{y})$ that enters into Eq. (30), which generically dilutes its contribution relative to that of $F(\mathbf{x})F(\mathbf{y})$ for the same values of the integrals of $|F(\mathbf{x})|^2$ and of $|\nabla G(\mathbf{x})|^2$ in Eq. (31), one can readily see that the maximum for X at fixed Y is obtained by setting $G(\mathbf{x}) = 0$, as well as choosing a spherically symmetric $F(\mathbf{x}) = f(|\mathbf{x}|/R)/R^2$ as given by Eq. (34). Then one gets that

$$\gamma = \frac{4\pi^2 \int_0^1 u^2 du f^2(u)}{\int_0^1 du \int_0^1 dv 2uv \ln \left| \frac{u+v}{u-v} \right| f(u)f(v)}. \quad (48)$$

One then sees that the minimum value for γ is

$$\gamma = \frac{4\pi^2}{\lambda}, \quad (49)$$

where λ is the largest eigenvalue of the weakly singular linear Fredholm integral equation of the third kind,

$$\int_0^1 dv 2 \ln \left| \frac{u+v}{u-v} \right| w(v) = \lambda w(u) \quad (50)$$

for $0 \leq u \leq 1$, with $w(u) = uf(u)$ being the eigenfunction.

For $F(\mathbf{x}) = f(|\mathbf{x}|/R)/R^2$ to be a smooth function of \mathbf{x} , f should be a smooth even function of its argument $u \equiv |\mathbf{x}|/R$. This means that $w(u)$ should be a smooth odd function of u , so we can expand it as an infinite sum of odd Legendre polynomials $P_{2m-1}(u)$,

$$w(u) = \sum_{m=1}^{\infty} c_m P_{2m-1}(u). \quad (51)$$

This expansion converts the integral eigenvalue Eq. (50) into the matrix eigenvalue equation

$$\sum_{n=1}^{\infty} A_{mn} c_n = \lambda \sum_{n=1}^{\infty} B_{mn} c_n, \quad (52)$$

where the matrix components are

$$\begin{aligned} A_{mn} &= \int_0^1 du \int_0^1 dv 2 \ln \left| \frac{u+v}{u-v} \right| P_{2m-1}(u) P_{2n-1}(v) \\ &= \frac{2}{[1 - 4(m-n)^2](m+n)(m+n-1)} \end{aligned} \quad (53)$$

and

$$B_{mn} = \int_0^1 du \int_0^1 dv P_{2m-1}(u) P_{2n-1}(v) = \frac{\delta_{mn}}{4m-1}. \quad (54)$$

[Actually, I cheated slightly in obtaining the explicit expression above for the matrix components A_{mn} . I calculated $A_{11} = 1$ by hand, but when I tried to calculate the general A_{mn} , I got finite sums that I did not readily see how to simplify. Therefore, I resorted to Maple. I did not quickly see how to get it to give me a simple general expression for A_{mn} either, but in one afternoon I was able to get it to give me all the values for $m < 10$, $n < 10$ (45 different terms, since $A_{mn} = A_{nm}$). The form of these terms was sufficiently simple that part way through their rather slow evaluation I was able to deduce the simple expression given in Eq. (53), which indeed fit all 45 terms. So although I have not bothered to find a rigorous proof that Eq. (53) is correct for all m and n not both smaller than 10, the fact that it is a very simple formula that works for all 45 smaller values strongly suggests that it is exact for all values of m and n . I could say that the proof is left as an exercise for the reader.]

Maple readily solved the matrix eigenvalue Eq. (52) for various truncations of the infinite matrices A_{mn} and B_{mn} . For example, 40-digit precision for 70×70 , 80×80 , 90×90 , 100×100 , 110×110 , and 200×200 truncations all gave the largest eigenvalue agreeing to 13 digits:

$$\lambda \approx 3.132010216749. \quad (55)$$

A 20-digit calculation of the 60×60 case gave the last digit 8 instead of 9 but was used to get the following approximate expansion of the eigenfunction corresponding to the largest eigenvalue:

$$\begin{aligned} w(u) \approx & + P_1(u) - 0.3968319408P_3(u) + 0.0102661635P_5(u) \\ & - 0.0070631137P_7(u) - 0.0032552106P_9(u) - 0.0018849293P_{11}(u) \\ & - 0.0011814233P_{13}(u) - 0.0007878354P_{15}(u) - 0.0005510316P_{17}(u) \\ & - 0.0004002345P_{19}(u) - 0.0002997366P_{21}(u) - 0.0002302203P_{23}(u) \\ & - 0.0001806208P_{25}(u) - 0.0001442924P_{27}(u) - 0.0001170806P_{19}(u) \\ & + \text{terms with coefficients less than } 0.0001. \end{aligned} \quad (56)$$

One can notice that only $P_1(u) = u$ and $P_3(u) = -1.5u + 2.5u^3$ give large contributions to the eigenfunction, so one can get a fairly accurate estimate of the largest eigenfunction by taking even just the 2×2 truncation of the matrices, which gives the eigenvalue

$$\lambda_2 = \frac{5(15 + \sqrt{57})}{36} \approx 3.131921449343, \quad (57)$$

which is smaller than the actual largest eigenvalue for the infinite matrices by less than one part in 35 283. An even simpler, but rather *ad hoc*, approximation is to

change the coefficient of the $P_3(x)$ term above to 0.4 and drop all the higher terms. Dividing this trial function for $w(u)$ by u (and multiplying by 8 to avoid fractions in the answer) gives $f(u) = 8 - 5u^2$, which may be inserted into Eqs. (48) and (49) to give another estimate for λ ,

$$\lambda_{\text{est}} = \frac{1757}{561} \approx 3.131907308378, \quad (58)$$

which has almost 16% more error than λ_2 , though this is still only a tiny error, being smaller than the actual largest eigenvalue for the infinite matrices by less than one part in 30 434. Even the very crude constant trial function for $f(u)$ gives an eigenvalue estimate, $\lambda_{\text{crude}} = 3$, that is smaller than the actual largest eigenvalue for the infinite matrices by only about 4.215%, or less than one part in 23.

Using the approximation Maple gave for λ , the largest eigenvalue of the infinite matrices, Eq. (49) then gives

$$\gamma = \frac{4\pi^2}{\lambda} \approx 12.604817632215, \quad (59)$$

which can be used in Eq. (47) to get the asymptotic behavior of $S_{\text{vN}}(x)$ at sufficiently small x . One can then see that this gives $S_{\text{vN}}(x) > x$ for

$$x < \gamma e^{1-\gamma} \approx 0.000115, \quad (60)$$

or alternatively for

$$S_{\text{vN}} < \gamma e^{1-\gamma} \approx 0.000115. \quad (61)$$

Thus if the dimensionless energy, $x \equiv 2\pi RE$, and the von Neumann entropy, $S = S_{\text{vN}} \equiv -\text{tr} \rho \ln \rho$, are sufficiently small, then with the definitions used here for these quantities, they can violate Bekenstein's conjectured entropy bound (1), $S \leq x$, though admittedly the range of x and S for which this happens is very narrow.

We may now use the value of γ , given by Eq. (59), in Eqs. (23), (45), and (46) to get $B \equiv S_{\text{vN}}/x$ as a precise implicit function purely of x , or, alternatively, to get both x and B as explicit functions of $a = \langle 0|h^2|0 \rangle$. Of course, this is merely for one simple example of a one-parameter family of mixed states given by Eq. (19), with q given by Eq. (44) and $|\psi\rangle$ given by Eq. (16) with $\beta = 0$ so $h = \alpha\chi$ and with χ given by Eq. (25) with $G(\mathbf{x}) = 0$ and Eq. (34) giving $F(\mathbf{x}) = f(|\mathbf{x}|/R)/R^2$ with $w(u) = uf(u)$ being an eigenvector corresponding to the largest eigenvalue, λ , of the homogeneous linear Fredholm integral equation (50). Therefore, it is not likely to give the maximum possible B as a function of x , which was called $B_{\text{N}}(x)$ in Eq. (9). However, this $B(x)$ does give at least a lower bound on $B_{\text{N}}(x)$ for a single massless scalar field.

5 Conjectures for Entropy Bounds of Vacuum-Outside- R States

I would conjecture that asymptotically at small $x \equiv 2\pi RE$, the density matrix (19), with all the entropy-maximization procedures given above for a density matrix of this form, gives $B(x)$ that does asymptotically approach the unknown global maximum function $B_N(x)$ for a single massless scalar field. Therefore, if we divide x into the asymptotic form of S_{vN} for small x that is given by Eq. (47), I would conjecture that this gives the asymptotic form of the true upper bound, $B_N(x)$, for very small x ,

$$B_N(x) \sim \frac{1}{\gamma} \left(\ln \frac{\gamma}{x} + 1 \right), \quad (62)$$

with Eq. (59) giving $\gamma \approx 12.604817632215$. One might expect a similar formula for other free massless fields, though perhaps each with a different value of γ .

When x is not small, it is certainly not the case that the density matrix of fixed x needed to maximize the entropy is approximately of the simple rank-two form given by Eq. (19). One would surely need a more general vacuum-outside- R state, with a density matrix obeying Eq. (2) (giving vacuum expectation values for all operators not in causal contact with the ball $r \leq R$ at $t = 0$), such as that given by Eq. (15), most likely with an infinite sum of terms and an infinite rank. I do not know how to proceed toward finding such a density matrix obeying Eq. (2) that would maximize $B(x) \equiv S/x$ at finite x that is neither asymptotically small or large. However, one might try using in Eq. (15) h_i 's that have the form given in Eq. (26), with the $F(\mathbf{x})$'s and $G(\mathbf{x})$'s of Eq. (25) being suitable eigenfunctions of the three-dimensional version of the integral equation (50). Even this wide class of examples may not be sufficient, since one could imagine instead constructing the Hermitian operators h_i from smeared functions of the field $\phi(t=0, \mathbf{x})$ and of its time-derivative (or conjugate momentum) $\dot{\phi}(t=0, \mathbf{x})$ that are not merely linear as is the χ given by Eq. (25). Going to nonlinear Hermitian operators (other than the relatively simple χ^2 considered above) leads to such a wealth of possibilities that I do not presently know how to proceed to obtain a true maximum for $B(x)$ at fixed finite x , the postulated function $B_N(x)$.

In the more usual case of boundary conditions on the field, quantum states obeying these boundary conditions may be coherently superposed (i.e., the corresponding wavefunctions added, not just the density matrices) to get other states that also obey the boundary conditions. Then one can look for superpositions that diagonalize the Hamiltonian (i.e., energy eigenstates). From these, one can form a Gibbs ensemble to maximize the von Neumann entropy at a fixed expectation value of the energy.

However, for the vacuum-outside- R states considered here, coherent superposi-

tions of pure vacuum-outside- R states are generically not vacuum-outside- R states. (Of course, positive-weight combinations of vacuum-outside- R density matrices are still vacuum-outside- R density matrices when normalized, since the vacuum-outside- R condition, that all expectation values outside the ball $r \leq R$ at $t = 0$ are the same as the vacuum, is homogeneous and linear in the density matrix, though not in the wavefunction.) Therefore, the procedure for diagonalizing the Hamiltonian for such states fails.

Indeed, one can see that none of the vacuum-outside- R states, except for the vacuum itself, can be an energy eigenstate. This is because any state which is non-vacuum in a finite region at some time will inevitably have that region spread with time. For fields with linear field equations, the perturbations of the field itself will spread. But even for fields with self-coupling which allow classical field solitons that do not spread with time, any quantum state of the field which is non-vacuum at some time will inevitably have that region spread with time as a result of the quantum uncertainty principle. For example, suppose that there is some definition of the location and momentum of the soliton, such that the velocity of the location is proportional to the momentum. Then the position-momentum uncertainty principle will prevent one from having that the location remain, with certainty, within any finite region for an infinite amount of time; the quantum uncertainty of the position, if initially confined to a finite region, will inevitably spread to extend all over space. Therefore, the confined configuration cannot be stationary and hence cannot be an energy eigenstate.

One can test my conjecture that Eq. (62) is the correct asymptotic form of the true upper bound on the entropy per $x \equiv 2\pi RE$ by examining some other simple density matrices of the form given by Eq. (15) that allow explicit evaluation of the von Neumann entropy. For example, the rank-three density matrix

$$\rho = (1 - q)|0\rangle\langle 0| + (q/2)e^{i\alpha x}|0\rangle\langle 0|e^{-i\alpha x} + (q/2)e^{-i\alpha x}|0\rangle\langle 0|e^{i\alpha x} \quad (63)$$

has $z \equiv x/\gamma = aq$, just like rank-two density matrix (19) when $\beta = 0$, and it has the three nonzero eigenvalues

$$\begin{aligned} p_1 &= \frac{1}{2}\left[1 - \frac{1}{2}q(1 - e^{-2a})\right] + \frac{1}{2}\sqrt{\left[1 - \frac{1}{2}q(1 - e^{-2a})\right]^2 - 2q(1 - q)(1 - e^{-a})^2} \\ &\approx 1 - q(1 - e^{-a}) + \frac{1}{2}q^2(1 - e^{-a})^2 \approx 1 - qa + \frac{1}{2}(q + q^2)a^2, \end{aligned} \quad (64)$$

$$\begin{aligned} p_2 &= \frac{1}{2}q(1 - e^{-2a}) \\ &\approx qa(1 - a), \end{aligned} \quad (65)$$

and

$$\begin{aligned}
p_3 &= \frac{1}{2}[1 - \frac{1}{2}q(1 - e^{-2a})] - \frac{1}{2}\sqrt{[1 - \frac{1}{2}q(1 - e^{-2a})]^2 - 2q(1 - q)(1 - e^{-a})^2} \\
&\approx \frac{1}{2}q(1 - q)(1 - e^{-a})^2 \approx \frac{1}{2}q(1 - q)a^2,
\end{aligned} \tag{66}$$

where the approximate equations apply for very small $a \equiv \alpha^2 X$. If one chooses q to maximize the von Neumann entropy of this mixed state,

$$S_{\text{vN}} \equiv -\text{tr}(\rho \ln \rho) = -p_1 \ln p_1 - p_2 \ln p_2 - p_3 \ln p_3, \tag{67}$$

for very small a (equivalently, very small z), one gets that $q \approx (1/3)(1 + 4a/3)$, so $z = aq \approx (a/3) + 4(a/3)^2$, which may be inverted to give $a \approx 3z(1 - 4z)$ and $q \approx (1/3)(1 + 4z)$. This then gives

$$S_{\text{vN}} \approx z[1 + z - (1 - z) \ln z] \sim \frac{x}{\gamma} \left(\ln \frac{\gamma}{x} + 1 \right), \tag{68}$$

which has the same asymptotic form for small x as Eq. (47) for the rank-two density matrix (19).

Another example would be to consider the rank-five density matrix

$$\begin{aligned}
\rho = (1 - q)|0\rangle\langle 0| &+ (q/4)e^{i\alpha\chi+i\beta\chi^2}|0\rangle\langle 0|e^{-i\alpha\chi-i\beta\chi^2} \\
&+ (q/4)e^{i\alpha\chi-i\beta\chi^2}|0\rangle\langle 0|e^{-i\alpha\chi+i\beta\chi^2} \\
&+ (q/4)e^{-i\alpha\chi+i\beta\chi^2}|0\rangle\langle 0|e^{i\alpha\chi-i\beta\chi^2} \\
&+ (q/4)e^{-i\alpha\chi-i\beta\chi^2}|0\rangle\langle 0|e^{i\alpha\chi+i\beta\chi^2},
\end{aligned} \tag{69}$$

which gives $z \equiv x/\gamma = q(a + b)$, where, as above, $a \equiv \alpha^2 X$ and $b \equiv 4\beta^2 X^2$. Even though the eigenvalue equation is now a fifth-order polynomial equation, it appears that one may be able to use the symmetries of the problem to find the eigenvectors and eigenvalues explicitly, as functions of q , a , and b , without requiring any roots higher than square roots. However, this seems to be messier than is worth doing here, so it shall be left as another exercise for the reader.

Nevertheless, one can show that when $a \ll 1$ and $b \ll 1$, there is one eigenvalue near unity, one near qa , one near $qb/2$, and the remaining two are smaller by factors of the order of $z = q(a + b)$. Therefore, in this limit only the three largest eigenvalues contribute significantly to the von Neumann entropy, giving

$$S_{\text{vN}} \approx -(1 - qa - qb/2) \ln(1 - qa - qb/2) - qa \ln(qa) - (qb/2) \ln(qb/2). \tag{70}$$

When this is maximized at fixed $z = q(a + b)$, one finds that the first three eigenvalues need to be approximately in a geometric series (as, e.g., are all the eigenvalues of

the thermal density matrix for an harmonic oscillator), giving $qb/2 \approx (qa)^2(1+qa)$. Solving for qa and qb in terms of z and inserting this back into Eq. (70) gives the maximum von Neumann entropy

$$S_{\text{vN}} \approx z[1 + \frac{1}{2}z - \ln z] \sim \frac{x}{\gamma} \left(\ln \frac{\gamma}{x} + 1 \right) \quad (71)$$

for the rank-five density matrix (69) at fixed tiny $z \equiv x/\gamma$. This entropy is just slightly larger, by an amount roughly $z^2[\ln(1/z) - (1/2)]$, than the corresponding maximum entropy (68) for the rank-three density matrix (63) at tiny $z = x/\gamma$, but it has the same asymptotic limit (given after the \sim sign).

Therefore, although of course these three simple examples of finite-rank density matrices do not begin to exhaust the infinite set of possibilities, they give some support to the conjecture given above that Eq. (62) is the correct asymptotic form of the upper bound for $S/(2\pi RE)$ when the denominator of this expression, $x \equiv 2\pi RE$, is much smaller than unity.

In the opposite limit, when $x \equiv 2\pi RE$ is much greater than unity, we would expect that, at least for a scale-invariant field so that the energy E is large with respect to all relevant parameters with the same dimension ($1/R$ being the only relevant one if the field does not have a rest mass or other parameter setting a higher energy scale), the maximum entropy is given by Eq. (11) for high-temperature thermal radiation, giving

$$B_{\text{N}}(x) \approx \frac{\beta}{x^{1/4}}, \quad (72)$$

with now

$$\beta = \left[\frac{2^6}{3^6 5} (n_b + \frac{7}{8} n_f) \right]^{\frac{1}{4}}, \quad (73)$$

no longer the β of Eq. (26) that we have subsequently set to zero to maximize $B(x)$ for the density matrix (19). It is tempting to combine this asymptotic formula for $x \gg 1$ with the asymptotic formula (62) conjectured above for $x \ll 1$ to conjecture that a reasonably good approximate formula for $B_{\text{N}}(x)$, as a function of any $x \equiv 2\pi RE$, for a quantum field theory with a given set of massless fields, is

$$B_{\text{N}}(x) \simeq \frac{4}{\gamma} \ln \left(1 + \frac{\beta\gamma}{4x^{1/4}} \right), \quad (74)$$

where the constants β and γ would depend upon the massless fields in the theory. For the single massless real scalar field that has been considered here, $n_b = 1$ and $n_f = 0$, so Eq. (26) would give

$$\beta = \left(\frac{2^6}{3^6 5} \right)^{\frac{1}{4}} \approx 0.364016115028, \quad (75)$$

and Eq. (59) has already given $\gamma \approx 12.604817632215$ for the single real massless scalar field.

Another way to state this conjecture is to write

$$B_N(x) = \frac{4}{\gamma} \ln \left(1 + \frac{\beta\gamma}{4x^{1/4}} \right) C(x), \quad (76)$$

where $C(x)$ is a correction factor yet to be found, and then conjecture that $C(x)$ tends asymptotically to unity for both very small and very large x , and perhaps further to conjecture that $C(x)$ is always relatively close to unity (e.g., say within a factor of two). This conjecture would then imply a conjectured entropy bound,

$$S_{\text{vN}} \leq \frac{8}{\gamma} \ln \left(1 + \frac{\beta\gamma}{4(2\pi RE)^{1/4}} \right) 2\pi RE. \quad (77)$$

An improved formula might be to write

$$B_N(x) = \frac{4}{\gamma} \ln \left(\frac{1 + Ax^{-1/4} + Bx^{-1/2}}{1 + Cx^{-1/4}} \right) \tilde{C}(x) \quad (78)$$

with $B/C = (e\gamma)^{1/4}$ to fit the final 1 in the asymptotic formula (62) for $x \ll 1$, $A - C = \beta\gamma/4$ to fit the asymptotic formula (72) for $x \gg 1$, and $2B - A^2 + C^2 = \gamma\delta/2$ to fit the following two-term improvement to Eq. (72):

$$B_N(x) \approx \frac{\beta}{x^{1/4}} + \frac{\delta}{x^{1/2}}. \quad (79)$$

I have not tried to work out what δ is. It would be straightforward to calculate, if it were the same as for the thermal state of a massless scalar field inside a sphere with Dirichlet boundary conditions on the field at the boundary $r = R$, but it is not obvious to me whether or not it is the same.

Since Eq. (78) with the correction factor $\tilde{C}(x)$ omitted should give a better asymptotic fit to $B_N(x)$ than Eq. (74), I would expect that $\tilde{C}(x)$ would generally be closer to unity than the corresponding correction factor $C(x)$ of Eq. (76) [not to be confused with the coefficient C in Eq. (78)]. But whether this is true over the entire infinite range of x remains to be seen.

For free massive quantum fields, for fixed entropy one would expect that the energy would have to be higher, so an upper bound for a set of free massless quantum fields should also give an upper bound for a corresponding set of free massive quantum fields. Therefore, I would conjecture that for any given free quantum field theory, one can find a β and γ (presumably with β obeying Eq. (75), and γ some combination of eigenvalues of the appropriate integral equations) such that the inequality (77) holds with that value of β and γ . The conjecture might even be true for any reasonable interacting quantum field theory that is causal in the sense defined above, though then one might need a different value of β .

6 Possibilities for Trying to Retain Bekenstein's Proposed Bound

Returning to a consideration of how Bekenstein's proposed bound (1) fits with the results derived and conjectured here, one first notes that the results here violate (1) for the von Neumann entropy S_{vN} of a vacuum-outside- R mixed state with sufficiently small energy expectation value E , e.g., for $x \equiv 2\pi RE < 0.000115$ in the example above. However, even if one accepts the use of vacuum-outside- R states for defining a finite size R , one might still object that the Bekenstein bound is not intended to be applied to the definition of E and/or S being used here.

For example, Schiffer and Bekenstein [8] refer to "quantum states accessible to the field system with energy up to and including E ." It could be objected that since the vacuum-outside- R states considered above are not energy eigenstates, they are actually composed of states with energy both lower and higher than the energy expectation value that I have used as the definition of E . If one takes E to be the energy of one of the energy eigenstates that is sufficiently higher than the expectation value, then the Bekenstein bound (1) may be obeyed even in the examples I have given above that violate the bound when E is taken to be the energy expectation value.

But if one takes this approach, it is hard to see how to give any content to the proposed bound for the vacuum-outside- R states. Presumably not only is it the case that any vacuum-outside- R state is not an energy eigenstate (since it is not stationary), but also it is surely the case that if any vacuum-outside- R state is decomposed into energy eigenstates, it will include energy eigenvalues of arbitrarily large value. However, using an arbitrarily large value of E in the bound (1) makes it trivial, entropy less than or equal to infinity. Therefore, for the bound to have any content, we need to have a definition of E that gives finite values. The definition given by Eq. (3) above, $E \equiv \text{tr}(H\rho)$, is surely the simplest, though others could be proposed.

For example, one could propose instead that for a vacuum-outside- R density matrix of the form (15), E could be defined as the maximum value of the expectation value of the Hamiltonian H in any of the normalized states $e^{ih_i}|0\rangle\langle 0|e^{-ih_i}$ whose sum, weighted by the q_i 's, forms ρ . However, using this definition would not avoid violations of Bekenstein's bound (1). For example, one could use the density matrix (63) with $q = 1$ so that the first term vanishes, and then the remaining two terms are of the form (15) with $h_1 = \alpha\chi$ and $h_2 = -\alpha\chi$, and with $q_1 = q_2 = 1/2$. Each of the two nonzero terms of the density matrix then gives the same energy expectation

value, and one can calculate that for small $z = x/\gamma = a \equiv \alpha^2 X$ one gets

$$S_{\text{vN}} \approx z[1 - \frac{1}{2}z - (1-z)\ln z] \sim \frac{x}{\gamma} \left(\ln \frac{\gamma}{x} + 1 \right), \quad (80)$$

again violating Bekenstein's bound (1) for $x < 0.000115$.

Yet another proposal would be that E be defined as E_{max} , the maximum expectation value of the Hamiltonian in all of the eigenstates of the density matrix. In the example just discussed,

$$\rho = p_1|1\rangle\langle 1| + p_2|2\rangle\langle 2| \quad (81)$$

with $p_1 = (1 + e^{-2a})/2$, $p_2 = (1 - e^{-2a})/2$, and orthonormal eigenstates

$$|1\rangle = \frac{\cos(\alpha\chi)|0\rangle}{\sqrt{p_1}} \quad (82)$$

and

$$|2\rangle = \frac{\sin(\alpha\chi)|0\rangle}{\sqrt{p_2}}. \quad (83)$$

Then by using Eq. (33), one can calculate not only that

$$\begin{aligned} \langle 0|e^{-ih_1}He^{ih_1}|0\rangle &\equiv \langle 0|e^{-i\alpha\chi}He^{i\alpha\chi}|0\rangle = \langle 0|e^{-ih_2}He^{ih_2}|0\rangle \equiv \langle 0|e^{i\alpha\chi}He^{-i\alpha\chi}|0\rangle \\ &= \alpha^2 Y \equiv \alpha^2 XZ \equiv aZ \end{aligned} \quad (84)$$

as given by Eq. (29), but also

$$\begin{aligned} \langle 0|e^{-ih_1}He^{-ih_2}|0\rangle &\equiv \langle 0|e^{-i\alpha\chi}He^{-i\alpha\chi}|0\rangle = \langle 0|e^{ih_2}He^{ih_1}|0\rangle \equiv \langle 0|e^{i\alpha\chi}He^{i\alpha\chi}|0\rangle \\ &= aZe^{-a}. \end{aligned} \quad (85)$$

From these results and from the form of the orthonormal eigenstates given by Eqs. (82) and (83), one readily obtains

$$H_{11} \equiv \langle 1|H|1\rangle = \frac{aZ(1 - e^{-a})}{1 + e^{-2a}} \approx \frac{1}{2}a^2Z \quad (86)$$

and

$$H_{22} \equiv \langle 2|H|2\rangle = \frac{aZ}{1 - e^{-a}} \approx Z \quad (87)$$

Then if one takes E_{max} , the larger of H_{11} and H_{22} , namely H_{22} , as the definition of E , one sees that it has the positive lower limit $Z = \gamma/(2\pi R)$ as one takes a (and hence S_{vN}) to zero, so with this definition one does not get a violation of Bekenstein's bound (1) in this example. In particular, for this example

$$\begin{aligned} B_{E_{\text{max}}} \equiv \frac{S_{\text{vN}}}{2\pi RE_{\text{max}}} &= \frac{1 - e^{-a}}{\gamma a} \left\{ -\frac{1}{2}(1 + e^{-2a}) \ln \left[\frac{1}{2}(1 + e^{-2a}) \right] \right. \\ &\quad \left. - \frac{1}{2}(1 - e^{-2a}) \ln \left[\frac{1}{2}(1 - e^{-2a}) \right] \right\} < 1. \end{aligned} \quad (88)$$

Whether this definition of $B_{E_{\max}}$ always gives a result less than unity for all vacuum-outside- R states, thus agreeing with Bekenstein's bound, remains to be proven, but the extremely meagre evidence that I have does not seem to contradict this conjecture.

Another objection that might be made against the violations of Bekenstein's bound (1) using vacuum-outside- R states to define R , using the expectation value of the Hamiltonian given by Eq. (3) as the definition of the energy E , and using the von Neumann entropy given by Eq. (4) as the definition of the entropy S , is to demand that the entropy instead be given by a microcanonical ensemble rather than by Eq. (4) applied to any mixed state. In other words, instead of allowing a generic vacuum-outside- R mixed state or density matrix ρ obeying Eq. (2) for all operators O completely confined to the region D , $r > R + |t|$, one might propose that Bekenstein's conjectured bound should only be applied to density matrices made up of equal mixtures of n orthogonal vacuum-outside- R pure states. (These are rank- n density matrices with precisely n nonzero eigenvalues, all equal to $1/n$, and with the corresponding set of n orthonormal eigenvectors all being vacuum-outside- R pure states.) The entropy of such a density matrix whose nontrivial part is proportional to the identity matrix in the n nontrivial dimensions is then $S = \ln n$.

Again, I do not have evidence that Bekenstein's conjectured bound (1) is violated for such restricted vacuum-outside- R density matrices. However, it is a rather severe limitation to restrict the discussion to such a small subset of vacuum-outside- R density matrices, a subset of measure zero in the space of all such density matrices. Furthermore, it appears rather difficult to find many explicit examples of precisely orthogonal vacuum-outside- R pure states.

For example, for each fixed choice of the two functions $F(\mathbf{x})$ and $G(\mathbf{x})$ in Eq. (25), the constants α and β in Eq. (26) give a two-parameter family of vacuum-outside- R pure states of the form $e^{i\alpha\chi + i\beta\chi^2}|0\rangle$, and then Eq. (27) and its trivial generalization gives the inner product between any two states among this two-parameter family. However, none of these inner products are zero for finite α 's and β 's (and hence for finite expectation values of the energy), so none of these states are orthogonal for fixed $F(\mathbf{x})$ and $G(\mathbf{x})$.

Of course, if one combines various ones of these nonorthogonal pure state density matrices to get a mixed density matrix and then finds the eigenvectors of that density matrix, they will form an orthonormal set of density matrices, such as the set $|1\rangle\langle 1|$ and $|2\rangle\langle 2|$ of Eqs. (82) and (83). However, these density matrices are not by themselves vacuum-outside- R states, but only when they are combined with the particular eigenvalues p_1 and p_2 given just before Eq. (82). Hence they cannot be used in a different linear combination (e.g., with $p_1 = p_2 = 1/2$) to get a vacuum-outside- R state that is an equal-weight combination of n orthonormal

pure-state vacuum-outside- R states.

The only explicit vacuum-outside- R state orthogonal to the vacuum itself (the trivial vacuum-outside- R state) that I have found so far is the pure state

$$|\psi\rangle = \exp(i\alpha e^{-\beta\chi^2})|0\rangle \quad (89)$$

with real α and β (not the same α and β used elsewhere in this paper) chosen so that

$$\langle 0|\psi\rangle = \sum_{n=0}^{\infty} \frac{(i\alpha)^n}{n} (1 + 2n\beta X)^{-1/2} = 0. \quad (90)$$

Using Maple, I found a numerical solution at

$$\alpha \approx 4.727048274, \quad \beta \approx 1.536994796/X. \quad (91)$$

I have not worked out the expectation value of the energy, $\langle \psi|H|\psi\rangle$, of this pure state, but I suspect that it is greater than $\ln 2/(\pi R)$, so that the entropy of an equal mixture of this state and of the vacuum, $\ln 2$, would be less than $2\pi R$ times the expectation value of the energy in this mixed state (half the expectation value of the energy of the pure state $|\psi\rangle\langle\psi|$, since the vacuum half of the mixed state contributes zero to the expectation value of the energy). If so, then this example would not be a counterexample to Bekenstein's conjectured bound (1) restricted to microcanonical ensembles that are equal mixtures of orthogonal pure vacuum-outside- R states.

It would be interesting to find the lowest-energy vacuum-outside- R state orthogonal to the vacuum itself and see whether its energy is indeed not more than $\ln 2/(\pi R)$, but I do not see how to do this at present. More generally, one would like to find, for each positive integer n , the set of n mutually orthogonal vacuum-outside- R states (possibly, but not necessarily, including the vacuum itself) such that the sum of the n energy expectation values, say E_s , is minimized. Then if one finds that $E_s \geq (n \ln n)/(2\pi R)$ for each n , then Bekenstein's conjectured bound (1) will be obeyed for these microcanonical ensembles of vacuum-outside- R states.

One way to look for other pure vacuum-outside- R states orthogonal to the vacuum would be to choose some non-hermitian operator, say κ , that is confined to the ball $r \leq R$ at $t = 0$, and consider the one-complex-parameter (two-real-parameter) set of states

$$|\psi(C; \kappa)\rangle = e^{iC\kappa - i\bar{C}\kappa^\dagger}|0\rangle \quad (92)$$

for the complex parameter C . For a generic such κ , $\langle 0|\psi(C; \kappa)\rangle$ would be a complex function of C (not analytic, since both C and its complex conjugate \bar{C} appear in the definition of $|\psi(C; \kappa)\rangle$), and a simple parameter-counting argument suggests that there should be discrete complex values of C at which $\langle 0|\psi(C; \kappa)\rangle =$

0, giving a pure vacuum-outside- R states orthogonal to the vacuum, though of course for particular κ 's, the number of such discrete solutions for C may be zero. One might extend this method to try to find n mutually orthogonal vacuum-outside- R states; this would require n different operators κ_i and $n(n+1)/2$ complex parameters. The obvious problem for carrying out this procedure explicitly is that for most sets of operators κ_i , the inner products (functions of the complex parameters) would be difficult to evaluate.

Perhaps a compromise to the stringent requirement of a microcanonical ensemble of n equally-weighted orthogonal pure vacuum-outside- R density matrices is simply to use the von Neumann entropy $S_{\text{vN}} \equiv -\text{tr} \rho \ln \rho$, which equals $\ln n$ for a microcanonical ensemble, but require that it be at least as large as $\ln 2$, the minimum nontrivial value for a microcanonical ensemble. Then one might conjecture that the bound (1) is correct for vacuum-outside- R states such that $S = S_{\text{vN}} \geq \ln 2$.

Alternatively, one might replace Bekenstein's conjectured bound (1) with the weaker conjectured bound

$$S \leq 2\pi ER + \ln 2, \quad (93)$$

still using $S = S_{\text{vN}} \equiv -\text{tr} \rho \ln \rho$, $E \equiv \text{tr}(H\rho)$, and restricting to vacuum-outside- R density matrices ρ obeying $\text{tr}(O\rho) = \langle 0|O|0 \rangle$ for all operators O totally confined to the region D , $r > R + |t|$, that is not in causal contact with the ball B , $r \leq R$ at $t = 0$ (no causal curves connecting these two regions). (Equivalently, one may require that $\text{tr}(O\rho) = \langle 0|O|0 \rangle$ for all operators O that commute with all operators defined totally on the ball B .)

7 Other Ways to Define a Radius R

So far I have been considering only the new proposal to define R by restricting to vacuum-outside- R states. However, one might ask whether there are other ways to define a radius R for a class of states for which one is seeking a bound on the entropy S as a function of R and of the energy E .

One proposal that is very close to my proposal of vacuum-outside- R states is a proposal for what might be called stressless-outside- R states, states such that on the $t = 0$ flat hyperplane of the Minkowski spacetime that I have always been assuming so far in this paper, the expectation value of the regularized stress-energy tensor operator, $T_{\mu\nu}$, is zero everywhere outside the radius R (as it is everywhere for the vacuum state),

$$\tau_{\mu\nu}(\mathbf{x}) \equiv \text{tr}(T_{\mu\nu}(t=0, \mathbf{x})\rho) = 0 \quad (94)$$

for all $|\mathbf{x}| > R$. Alternatively, one might restrict to what might be called energyless-

outside- R states,

$$\varepsilon(\mathbf{x}) \equiv \tau_{00} \equiv \text{tr}(T_{00}(t=0, \mathbf{x})\rho) = 0 \quad (95)$$

for all $|\mathbf{x}| > R$, for which the expectation value $\varepsilon(\mathbf{x})$ of the regularized energy density operator, T_{00} , at $t = 0$ vanishes outside the radius $|\mathbf{x}| = R$.

Of course, all vacuum-outside- R states are also stressless-outside- R states, and all stressless-outside- R states are also energyless-outside- R states, but I do not know whether the converses of these statements are true. If they are not both true, there would exist energyless-outside- R states, and possibly also stressless-outside- R states, that are not also vacuum-outside- R states. If there is indeed a broader class of states than vacuum-outside- R states, whether stressless-outside- R states and/or energyless-outside- R states, then the corresponding entropy maximization function $\sigma_{\text{vN}}(R, E)$ would be expected to be larger for the broader class of states.

One can try to define R for even broader classes of states, not by requiring that some expectation values vanish for $r > R$ at $t = 0$, but instead by using the spatial distribution of some quantity to define an effective radius R . One way that first comes to mind is to use some spatially-dependent real weight function $W(\mathbf{x})$ coming from the quantum state to define R as an rms value of $r \equiv |\mathbf{x}|$:

$$R_W^2 = \frac{\int d^3x W(\mathbf{x}) r^2}{\int d^3x W(\mathbf{x})}. \quad (96)$$

Of course, the weight function should be such that both the numerator and the denominator are finite and have the same sign (which without loss of generality will be assumed to be positive), at least for the class of states to be considered.

An obvious simple choice of the weight function is the energy density expectation value $\varepsilon(\mathbf{x})$. Then the denominator of Eq. (96) for R_W^2 is the total energy, which is positive for a nontrivial state. However, the numerator is not positive for all nontrivial states, as one can see from the following argument: Motivated by the state with locally negative energy density given by Kuo and Ford [37], consider the state

$$|\psi\rangle = \alpha|0\rangle + \beta|2\rangle, \quad (97)$$

where $|\alpha|^2 + |\beta|^2 = 1$ and $|2\rangle$ is a two-quantum state of energy

$$E_2 \equiv \langle 2|H|2\rangle = \int d^3x \langle 2|T_{00}|2\rangle \quad (98)$$

and with mode functions that are sufficiently localized that

$$R_2^2 E_2 \equiv \int d^3x \langle 2|T_{00}|2\rangle r^2 \quad (99)$$

is finite. Because T_{00} is the regularized (e.g., normal-ordered) energy density operator, $\langle 0|T_{00}|0\rangle = 0$ and $\int d^3x \langle 2|T_{00}|0\rangle = \int d^3x \langle 0|T_{00}|2\rangle = 0$. For a generic

two-quantum state $|2\rangle$,

$$C \equiv \int d^3x \langle 2|T_{00}|0\rangle r^2 \quad (100)$$

will be a nonzero complex number. Then Eq. (96) for $W(\mathbf{x}) = \varepsilon(\mathbf{x})$ gives

$$R_W^2 = \frac{\int d^3x \langle \psi|T_{00}|\psi\rangle r^2}{\int d^3x \langle \psi|T_{00}|\psi\rangle r^2} = R_2^2 + \Re\left(\frac{2C}{E_2} \frac{\alpha}{\beta}\right). \quad (101)$$

Since α/β can be an arbitrary complex number, R_W^2 can take any real value if $C \neq 0$, including zero and negative values. Even if one restricted to states for which R_W^2 is positive, this quantity can be made arbitrarily small, and then if such a state has finite energy and positive entropy (e.g., by being a mixture of the vacuum state $|0\rangle\langle 0|$ and of $|\psi\rangle\langle\psi|$), it can make $B \equiv S/(2\pi RE)$ arbitrarily large.

If one did want to use $\varepsilon(\mathbf{x})$ as the weight function $W(\mathbf{x})$ in Eq. (96), one would have to restrict the states so that R_W^2 cannot be too small for states of finite energy and nonzero entropy. One way that might work would be to restrict the states to those in which $\varepsilon(\mathbf{x})$ is nonnegative everywhere, unlike the state $|\psi\rangle\langle\psi|$ for sufficiently large $-C\alpha/\beta$.

Another option that might work for all sufficiently localized states would be to choose a weight function $W(\mathbf{x})$ that is nonnegative for all states. Examples of this would be ε^2 , $\sum_{\mu=0}^4 \sum_{\nu=0}^4 (\tau_{\mu\nu})^2$, $(\tau_\mu^\mu)^2$, $(\tau_\nu^\nu \tau_\mu^\nu)^2$, $(\tau_\nu^\mu \tau_\rho^\nu \tau_\mu^\rho)^2$, $(\tau_\nu^\mu \tau_\rho^\nu \tau_\sigma^\rho \tau_\mu^\sigma)^2$, $(\text{tr}(\phi(t=0, \mathbf{x})\rho))^2$, $(\text{tr}(\dot{\phi}(t=0, \mathbf{x})\rho))^2$, $(\text{tr}(\phi^2(t=0, \mathbf{x}) : \rho))^2$, $(\text{tr}(\dot{\phi}^2(t=0, \mathbf{x}) : \rho))^2$, $(\text{tr}(\phi^{;\mu}(t=0, \mathbf{x})\phi_{;\mu}(t=0, \mathbf{x}) : \rho))^2$, etc., and positive powers of these positive functions of \mathbf{x} at $t=0$.

The fourth quantity above is the square of $\tau_\nu^\mu \tau_\mu^\nu$, which itself usually seems to be positive everywhere, though it might be some restriction of states for this to be true everywhere for all states in the class.

If one used one of these positive quantities, or one of the usually positive quantities just for states in which it is everywhere positive, as a weight function $W(\mathbf{x})$ for defining R by Eq. (96), one would again presumably get some upper bound on the entropy S as a function of R and of E (depending on how one defined S and E and what further restrictions one puts on the states). However, I have not investigated what these relations might be.

8 Difficulties with Entropy Bounds in Quantum and Semiclassical Gravity

So far I have been restricting attention to nongravitational quantum field theories in flat Minkowski spacetime. However, since the original motivation for Bekenstein's conjectured entropy bound came from quantum considerations of gravitating black

holes, it is interesting to consider whether a similar entropy bound can be applied in quantum gravity.

Here I must admit that I see serious problems in attempting to apply the bound to quantum gravity. Assuming that the quantum part of quantum gravity is sufficiently similar to the ordinary quantum theory of nongravitational systems, the entropy S might still be a well-defined quantity, at least for a complete system, such as the entire universe (though if the ultimate quantum gravity theory specifies a unique quantum state, there may be no option as to what the entropy is). The energy E is more problematic, at least if the universe is not asymptotically flat, though if one can restrict to quantum states in which the universe is asymptotically flat, then E also might have a good definition in quantum gravity. However, what I don't see how to give a good precise definition for is the size R .

The main problem is that states of quantum gravity should be coordinate invariant, so it is hard to see how to say that some state is confined to a radius R or has this size. How would one define the center with respect to which the state is within a distance R ? Furthermore, if the state is an asymptotically flat one with energy E , the gravitational field of this energy should extend all the way out to spatial infinity, so in that sense it seems that the state cannot be confined to be within radius R or have vacuum properties outside that radius.

It is not that I have a rigorous proof that an entropy bound such as Eq. (1) cannot be applied in quantum gravity, but I just don't see how it can be applied.

The situation seems somewhat more hopeful in semiclassical gravity, in which one has quantum field theory for nongravitational fields on a classical curved spacetime (perhaps whose Einstein tensor is proportional to the regularized expectation value of the stress-energy tensor of the nongravitational quantum fields). In this case one can imagine defining the equivalent of vacuum-outside- R states of energy E in the following way:

Take an asymptotically flat spherically symmetric spacetime which has a totally geodesic Cauchy hypersurface (with zero extrinsic curvature, say at $t = 0$), about which it has time-reflection symmetry ($t \rightarrow -t$). Use a Schwarzschild radial coordinate r , the circumference/ (2π) of each symmetrical sphere. Outside the $r = R$ sphere on this $t = 0$ hypersurface, assume that the spatial metric and the expectation value of all operators confined to this region are the same as that of the static spherically symmetric asymptotically Schwarzschild semiclassical metric with ADM mass E and quantum state that is the semiclassical version of the zero-temperature Boulware state for this metric. The entropy S can then be the von Neumann entropy $S_{\text{vN}} \equiv -\text{tr} \rho \ln \rho$ of the quantum state of the nongravitational quantum field in the classical curved metric.

If one applies this definition to all possible states of this form, then it seems

one can easily violate a bound of the form (1) by a state with arbitrarily large entropy by having the $r = R$ sphere be a neck separating the asymptotically flat exterior with an interior on the $t = 0$ hypersurface that is almost an entire three-sphere of arbitrarily large size filled with thermal radiation. In other words, take the interior of the $r = R$ two-sphere to be the moment of maximum expansion of an almost-complete large $k = 1$ radiation Friedman-Robertson-Walker model, and take the exterior to be a moment of time-symmetry of a nearly empty approximately Schwarzschild metric. If the interior three-sphere radius is $a \gg R$, giving an interior volume going as a^3 , the semiclassical Einstein equations imply that the radiation energy density must go as a^{-2} in Planck units at the moment of maximum expansion, so the temperature T goes as $a^{-1/2}$ and the entropy density goes as T^3 or as $a^{-3/2}$. When this is multiplied by the volume, one gets an entropy going as $a^{3/2}$, which can be made arbitrarily large for fixed R and E (the asymptotic ADM mass) by making the interior size a arbitrarily large.

One might seek to avoid this violation of (1) by restricting the states to which it is conjectured to apply to exclude this example of a huge interior universe separated from an asymptotically flat exterior by a relatively small neck.

One way to do that would be to demand that inside the $r = R$ two-sphere on the hypersurface of time symmetry, there are no round two-spheres of radius greater than R (i.e., topological two-spheres with intrinsic two-metrics that are those of the standard unit round two-sphere multiplied by a constant r^2 that is larger than R^2). This would exclude an interior that is approximately a large round three-sphere of radius $a \gg R$, since such an interior region would have round two-spheres of radii $r \approx a$. However, it still appears to allow a very long throat of radius near R , which by its arbitrarily great length could have arbitrarily large volume and hence arbitrarily large entropy S for fixed R and E .

Another way to restrict the states so that they might possibly obey a bound similar to (1) is to demand that the evolution of the semiclassical geometry give a nonsingular metric over the whole of \mathbf{R}^4 . This would exclude the examples of a large internal approximate three-sphere and also the long internal throat, since these examples would be expected to collapse gravitationally to singularities. Only in cases in which the metric is not too much different from flat spacetime would one expect that no singularities develop from gravitational collapse, and in these cases one might expect an entropy bound not too different from its flat spacetime form. However, for the restricted states to be sufficiently broad to encompass most of the semiclassical gravity generalizations of the allowed states (e.g., vacuum-outside- R states) in the nongravitational theory, the semiclassical Einstein equations should give nonsingular evolution in these cases of sufficiently weak gravity. Since the semiclassical Einstein equations are of higher order in time than the ordinary Einstein equations with a

classical source, it is not clear that this will be the case, and so one might need to find a suitable semiclassical gravity theory before trying to apply entropy bounds.

9 Conclusions and Acknowledgments

In conclusion, we have found that one can formulate precise definitions for entropy bounds of a complete quantum field system (i.e., one not restricted to the interior of some boundary) by giving precise definitions for the size R of the system, at least when the metric is classical so that sizes can be unambiguously defined. In particular, R may be defined for vacuum-outside- R states as the largest round two-sphere on a suitable $t = 0$ hypersurface, outside of which all of the operators have the same expectation values as in a suitable vacuum state (e.g., the ordinary vacuum state for nongravitational fields in Minkowski spacetime, or a Boulware-type quantum state in semiclassical gravity). Other values of R may also be defined, such as the rms value of r with a suitable weight function dependent upon the quantum state of the field. On the other hand, for a fully quantum gravity theory, it appears to be difficult to give an unambiguous definition of a size R of a system, so it is not clear there how to define a bound for the entropy S in terms of the energy E and a size R .

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